

FLUIDS IN THE CSA Cu-Pb-Zn DEPOSIT, NSW, AUSTRALIA

GILES, A. D. AND MARSHALL, B. THE UNIVERSITY OF TECHNOLOGY, SYDNEY, AUSTRALIA OF TECHNOLOGY, SYDNEY, AUSTRALIA.

CSA is a Cobar-type deposit. Cobar-type Au-Cu-Ag-Pb-Zn massive-sulfide ores occupy epigenetic systems that have dip, strike and width extents of 1 km, 300 m and 100 m respectively. The type-model involves structurally-controlled emplacement during basin-inversion tectonism spanning late diagenesis and regional metamorphism. Sulfides deposited from two fluid sources during the variable and progressive overprinting of saline, CH₄-rich, basinal fluids by higher temperature, low-salinity, H₂O-CO₂ bearing, basement-derived, metamorphic(?) fluids; or simply by these fluids mixing. Ag-Pb-Zn mineralization is said to be basin-dominated; Cu-Au is basement-dominated.

Five categories of secondary fluid inclusion (Types IA, IB, II, III, IVA) exist in quartz coexisting with sulfides in CSA ore. Primary fluid inclusions (Type IVB) occur in quartz crystals in vugs in syn- or post-cleavage tension gashes cross-cutting ore. In formation order (oldest first): Types IA (along microfractures in strained quartz) and IB (along grain boundaries in recrystallized quartz) are decrepitated and empty or have become infilled by later CH₄-rich fluid; Type II monophasic inclusions are CH₄ liquid (Th(CH₄) -90.1 to -84.7°C); Type III H₂O+CH₄ inclusions comprise CH₄ liquid immiscible within H₂O (Th(CH₄) -128.6 to -86.1°C, T_m(ice) -2.0 to -0.1°C, Th(total) 218 to 226°C); Type IVA are hydrosaline with traces of CH₄ gas (T_m(ice) -2.0 to -0.6°C, Th(to liquid) 113 to 305°C); and Type IVB growth-zone primaries are hydrosaline (T_m(ice) -1.5 to -0.9°C, Th(to liquid) 165.8 to 185.6°C). All of the inclusion-types are syn- to post-metamorphic; none provides direct evidence of late diagenetic connate fluid; none contain detectable CO₂; the original contents of Types IA and IB were lost through decrepitation. The populations suggest post-decrepitative outflow of CH₄-rich fluids during the syn- to post-metamorphic periods, coupled with progressive involvement of low-salinity, H₂O-rich meteoric(?) fluid. The results are indirectly consistent with basinal CH₄ liquids being involved with ore formation, but provide no insight on a higher temperature H₂O-CO₂ basement-sourced fluid.